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NONLINEAR EQUATORIAL SPREAD F: SPATIALLY LARGE BUBBLES RESULTIN--ETC(U)

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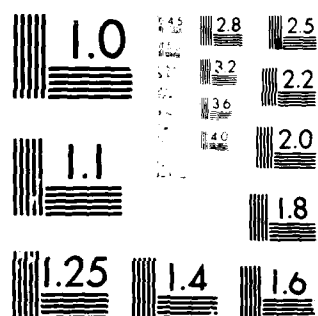
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NRL Memorandum Report 4154

**Nonlinear Equatorial Spread F: Spatially Large  
Bubbles Resulting from Large Horizontal Scale  
Initial Perturbations**

S. T. ZALESK AND S. L. OSSAKOW

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## 20. Abstract (Continued)

scale length perturbations previously used. Further, we find the level of plasma depletion inside the large scale bubbles to be appreciably higher than that of the smaller scale bubbles, approaching 100%, in substantial agreement with the observations. This level of depletion is due to the fact that the plasma comprising the large scale bubbles has its origin at much lower altitudes than that comprising the smaller scale bubbles. Analysis of the polarization electric fields produced by the vertically aligned ionospheric irregularities show this effect to be due to fringe fields similar in structure to those produced at the edge of a parallel plate capacitor.

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## I. INTRODUCTION

Our previous numerical simulations of the nonlinear evolution of the collisional Rayleigh-Taylor instability in the nighttime equatorial ionosphere (spread F) were confined to small ( $\sim 3\text{km}$ ) horizontal scale length initial perturbations and hence to fully developed spread F "bubbles" of approximately the same size in horizontal extent [Scannapieco and Ossakow, 1976; Ossakow et al., 1979], although spatially large vertically. However, observations by McClure et al [1977] also indicate ionospheric ion density "biteouts" of much larger horizontal extent ( $10 - > 200\text{ km}$ ) and greater intensity (ion density depletions up to three orders of magnitude) than indicated by our small scale simulations. Therefore, we have extended our previous calculations and have performed a series of numerical simulations using much larger horizontal length scales ( $\sim 75\text{km}$ ) for our initial perturbations. These seed long wavelength perturbations, for example, could be due to neutral atmosphere gravity wave effects [Rottger, 1976; Klostermeyer, 1978; Booker, 1979]. At the same time we have made very substantial improvements in the numerical techniques used to perform the simulations, including the utilization of the recently developed fully multidimensional flux-corrected transport (FCT) techniques of Zalesak [1979]. The results of our simulations indicate the following:

- 1) large horizontal scale length initial perturbations evolve nonlinearly into large horizontal scale length equatorial spread F "bubbles;"
- 2) these bubbles evolve on approximately the same time scale as do their smaller horizontal scale length counterparts;

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- 3) the plasma comprising these bubbles has its origin at much lower altitudes than that of the smaller horizontal scale length bubbles, resulting in plasma density depletions of very close to 100%.

This last result is due to the fact that the polarization fields produced by ionospheric irregularities, aligned vertically, possess a fringe field component whose vertical extent is proportional to the horizontal extent of the irregularities producing the field. This is quite similar in origin to the fringe field produced at the edge of a parallel plate capacitor. Since the vertical extent of this fringe field determines the minimum altitude from which the rising bubble can draw plasma, it is not surprising that larger horizontal scale bubbles are more severely depleted. In Section II we give a brief review of the relevant theory, and of the basic equations used in our simulations. Section III contains the numerical simulations and a physical interpretation of the results is given. Section IV contains a summary, and in the appendix we describe briefly the numerical techniques used in our present computer code, emphasizing the differences between the present code and our previous one.



## II. THEORY

The geometry of the physical problem we are modeling is the same as in Ossakow et al. [1979]. All our simulations are carried out in a two dimensional (x,y) coordinate system. The constant magnetic field  $\underline{B}$  is aligned along the  $\hat{z}$  axis (pointing north). Gravity is directed along the negative  $\hat{y}$  axis. Since our equations are two dimensional,  $n(y)$ ,  $\nu_R(y)$ , and  $\nu_{in}(y)$  are taken to be representative values of the ambient electron density, recombination coefficient, and ion-neutral collision frequency (the result of integrating these quantities along magnetic field lines and dividing by a normalizing scale length). Magnetic field lines are assumed to terminate at both ends in an electrically insulated medium (currents must close in the two dimensional plane, not in some distant E region).

Following Ossakow et al. [1979], we describe the system with the two-fluid plasma continuity and momentum equations:

$$\frac{\partial n_\alpha}{\partial t} + \nabla \cdot (n_\alpha \underline{v}_\alpha) = - \nu_R n_\alpha \quad (1)$$

$$\left( \frac{\partial}{\partial t} + \underline{v}_\alpha \cdot \nabla \right) \underline{v}_\alpha = \frac{q_\alpha}{m_\alpha} \left( \underline{E} + \frac{\underline{v}_\alpha \times \underline{B}}{c} \right) + \underline{g} - \nu_{\alpha n} (\underline{v}_\alpha - \underline{U}_n) \quad (2)$$

where the subscript  $\alpha$  denotes the species (i for ions, e for electrons),  $n$  is the species number density,  $\underline{v}$  is velocity,  $\nu_R$  is the recombination coefficient,  $\underline{E}$  is the electric field,  $\underline{g}$  is the gravitational acceleration,  $q$  is the species charge,  $\nu_{\alpha n}$  is the species collision frequency with the

neutral atmosphere,  $\underline{U}_n$  is the neutral wind velocity,  $c$  is the speed of light, and  $m$  is the species mass.

Note that, in contrast to Ossakow et al [1979], we have dropped the term  $+ \nu_R n_{\alpha 0}$  from (1). This is the equivalent of dropping the assumption of the existence of an ionization source given by that term. This ionization source was such that the ambient ionization profile  $n_{\alpha 0}(y)$  was an equilibrium profile ( $\partial n_{\alpha 0} / \partial t = 0$ ). Our present model therefore has instead

$$\frac{\partial n_{\alpha 0}}{\partial t} = - \nu_R n_{\alpha 0} \quad (3)$$

Hence, when normalized results  $n_{\alpha}(x,y)/n_{\alpha 0}(y)$  are later presented, it should be understood that both the numerator and denominator are time dependent.

Figure 1 shows the recombination rate  $\nu_R$  and ion-neutral collision frequency  $\nu_{in}$  used in our simulations. It shall be seen presently that  $\nu_{en}$  need not be specified as long as it is much smaller than the electron gyro frequency  $\Omega_e$ . For details on the form of  $\nu_{in}$  and  $\nu_R$  as depicted in Figure 1, see Ossakow et al.[1979]. If we neglect the inertial terms (the left-hand side) of (2) by assuming the inertial time scales are much larger than either the gyro periods or the mean time between collisions, then the equation can be inverted to give an algebraic expressions for  $\underline{v}_{\alpha}$ . In two-dimensional  $(x,y)$  geometry with  $\underline{B}$  along the  $\hat{z}$  axis, the solution is for our problem, with  $m_e \ll m_i$ ,  $\nu_{in}/\Omega_i \ll 1$ ,  $\nu_{en}/\Omega_e = 0$  (where  $\Omega_{\alpha} = eB/m_{\alpha}c$ ), and  $\underline{U}_n = 0$ .

$$\underline{v}_e = \frac{c}{B} \underline{E} \times \hat{z}, \quad \hat{z} = \frac{B}{|B|} \quad (4)$$

$$\underline{v}_i = \left( \frac{g}{\Omega_i} + \frac{c}{B} \underline{E} \right) \times \hat{z} + \frac{v_{in}}{\Omega_i} \left( \frac{g}{\Omega_i} + \frac{c}{B} \underline{E} \right) \quad (5)$$

We now make the electrostatic approximation,

$$\underline{E} = \nabla_{\perp} \phi \quad (6)$$

where  $\nabla_{\perp} \equiv \hat{x}(\partial/\partial x) + \hat{y}(\partial/\partial y)$ , and the quasi-neutrality approximation  $n_e \approx n_i \equiv n$ . We then have

$$\nabla_{\perp} \cdot \underline{j} = 0 \quad (7)$$

$$\underline{j} \equiv en (\underline{v}_i - \underline{v}_e) \quad (8)$$

Substituting (4) and (5) into (8) and evaluating (7), we have for the electrostatic potential:

$$\nabla_{\perp} \cdot (v_{in} n \nabla_{\perp} \phi) = - \frac{m_i}{e} g \frac{\partial}{\partial y} (v_{in} n) - \frac{B}{c} g \frac{\partial n}{\partial x} \quad (9)$$

As in Ossakow et al. [1979] we set  $\phi = \phi_0 + \phi_1$  where  $\nabla_{\perp} \phi_0 = - (m_i g / e) \hat{y}$ . Since  $\nabla_{\perp}^2 \phi_0 = 0$ , our final potential equation becomes

$$\nabla_{\perp} \cdot (v_{in} n \nabla_{\perp} \phi_1) = - \frac{Bg}{c} \frac{\partial n}{\partial x} \quad (10)$$

The effect of  $\phi_0$  is merely to superimpose a bulk westward plasma velocity  $g/\Omega_i$  on the electron velocity field determined from  $\phi_1$ , without affecting the morphology of the developing structures. Hence, we ignore this motion.

Our assumption of quasi-neutrality has made one of our two continuity equations (1) redundant. We therefore choose the electron equation for its simplicity:

$$\frac{\partial n}{\partial t} - \frac{\partial}{\partial x} \left( \frac{nc}{B} \frac{\partial \phi_1}{\partial y} \right) + \frac{\partial}{\partial y} \left( \frac{nc}{B} \frac{\partial \phi_1}{\partial x} \right) = -v_R n \quad (11)$$

### III. NUMERICAL SIMULATION RESULTS AND DISCUSSION

Equations (10) and (11), together with appropriate boundary conditions, constitute the nonlinear system of equations we shall solve numerically. Note that in contrast to Ossakow et al. [1979], we have chosen not to put the equations into a normalized form by dividing through by  $n_0(y)$ . The numerical techniques used to solve these equations are discussed in the appendix.

The numerical calculations to be presented were performed on a two-dimensional cartesian (x,y) mesh using 42 points in the x (east-west) direction, and 142 points in the y (vertical) direction. The (uniform) grid spacing was 2km in the y direction for all calculations. The grid spacing in the x direction was 200m in the "small" horizontal scale length cases and 5km in the "large" cases. The bottom of the grid corresponds to 252km altitude and the top of the grid to 534km altitude in all simulations. Periodic boundary con-

ditions were imposed on both  $n$  and  $\phi_1$  in the  $x$ -direction. In the  $y$  direction transmissive boundary conditions were imposed on  $n(\partial n / \partial y = 0)$  and Neumann ( $\partial \phi_1 / \partial y = 0$ ) boundary conditions were imposed on  $\phi_1$ .

Three kinds of plots will be presented: (1) contours of constant  $n(x,y,t)$ ; (2) contours of constant  $n(x,y,t)/n_0(y,t)$ ; and (3) contours of constant electrostatic potential  $\phi_1$ . Superimposed on each contour plot is a dashed line depicting  $n_0(y,t)$  for reference purposes. Our  $n_0(y,0)$  profile is such that the F2 peak is located at 434km altitude, and the minimum electron density scale length  $L = n_0 (\partial n_0 / \partial y)^{-1}$  is 10km. The simulations are all identical except for the east-west grid spacing  $\Delta x$  and the form of the initial perturbation. Two kinds of initial perturbations were used:

$$\text{Perturbation A: } \left\{ \begin{array}{l} \frac{n(x,y,0)}{n_0(y,0)} = 1 - e^{-3} \cos\left(\frac{\pi x}{8\Delta x}\right), \quad 0 \leq |x| \leq 8\Delta x \\ \frac{n(x,y,0)}{n_0(y,0)} = 1 - e^{-3} \frac{1}{2} \left[ \cos\left(\frac{\pi x}{8\Delta x}\right) - 1 \right], \quad 8\Delta x \leq |x| \leq 16\Delta x \\ \frac{n(x,y,0)}{n_0(y,0)} = 1, \quad |x| > 16\Delta x \end{array} \right. \quad (12)$$

$$\text{Perturbation B: } \frac{n(x,y,0)}{n_0(y,0)} = 1 - e^{-3} \cos\left(\frac{\pi x}{20\Delta x}\right) \quad (13)$$

Perturbation A is exactly the form used in Ossakow et al. [1979], and perturbation B is a pure sine wave of wavelength  $40 \Delta x$  (our system length in the x direction). Both represent maximum initial perturbation amplitudes of approximately 5%. Four simulations have been run: (i) 1S-Perturbation A with  $\Delta x = 200\text{m}$ ; (ii) 1L-Perturbation A with  $\Delta x = 5\text{km}$ ; (iii) 2S-Perturbation B with  $\Delta x = 200 \text{ m}$ ; and (iv) 2L-Perturbation B with  $\Delta x = 5\text{km}$ . Calculation 1S above is identical to ESF III of Ossakow et al. [1979]. The "large" versus "small" comparison obviously involves comparing calculation 1S to calculation 1L, and calculation 2S to calculation 2L. One notes that for the minimum  $L \approx 10\text{km}$  in our simulation,  $kL > 1$  for the 1S and 2S cases and  $kL < 1$  for the 1L and 2L cases.

Figure 2 shows isodensity contours of calculation 1S at times 300, 700, 1000, and 1200 seconds after initialization. Figure 3 shows the same contours at the same times but for calculation 1L. The presence of much lower density fluid in the bubble in calculation 1L is obvious. Also obvious is a basic difference in the bubble morphology at late times. At 1200 seconds, 1S has pinched off into two bubbles, with the more intense one below the initial central bubble. In addition, another bubble has formed in the sides of the calculation. These structures are more obvious in the plot of  $n(x,y)/n_0(y)$  at 1200 seconds for 1S shown in Fig. 4a. The maximum depletion levels are 70% in the top central bubble, 97% in the lower central bubble, and 95% in the side bubble. Note that here and in all subsequent plots of  $(n,x,z)/n_0(z)$  the contour plotting is such that the first

(outer) depletion contour  $n/n_0$  is 0.5 and each succeeding inner contour is 0.5 times the previous one. For example, the lower bubble in Fig. 4a has five contours. The outermost would have  $n/n_0 = 0.50$  (50% depletion), the next inner one  $n/n_0 = 0.25$  (75% depletion), the second inner one  $n/n_0 = 0.125$  (87.5% depletion) and the innermost  $n/n_0 = 0.03$  (97% depletion). For the enhancement contours (dashed lines) the first outer contour is 2.0 and the succeeding inner ones are 2.0 times the previous ones. In obtaining percentage enhancements and depletions one then subtracts 1.0.

Calculation 1L, on the other hand, shows a single plume of depleted ionization at 1200 seconds, with no secondary central bubble and no side bubble. There also is no indication of a widening of the top of the bubble, as there is in 1S. In Fig. 4b we show a plot of  $n(x,y)/n_0(y)$  for 1L at 1200 sec. The level of depletion is greater than 99.9% for the entire 10km by 70km oval "hole" located inside the tenth solid contour of Fig. 4b and represents at least a three order of magnitude decrease (biteout) in plasma density.

We now repeat the above comparison, but this time for perturbation B (calculations 2S and 2L). Figure 5 shows isodensity contours of calculation 2S at times 300, 700, 1000, and 1091 seconds after initialization; while Figure 6 shows similar plots of calculation 2L at times 700, 1000, 1200 and 1364 sec. Comparison again shows the presence of much lower density plasma in the bubble in the 2L calculation. Morphological differences are also present, the most notable being the widening of the top of the bubble in 2S

which is not present in 2L. Figure 7 shows a comparison of the  $n/n_0$  profiles at late time. Again maximum depletions in the 2S case are about 97%, while a large portion of the 2L plume is 99.9% depleted or greater.

We can also compare the effect of the form of the perturbation by comparing 1S to 2S and 1L to 2L. The latter comparison shows striking similarity, whereas the former shows some differences, the most notable being the lack of central bubble "pinching" and the lack of lateral bubbles in case 2S. We conclude that the morphology of the late-time bubbles is dependent, at least somewhat, on the form of the initial perturbation.

Bubble rise velocities are of some interest, and we give below the average bubble rise velocity for each case, computed from the last two frames of Figs. 2, 3, 5, and 6:

1S	210 m/sec
1L	230 m/sec
2S	420 m/sec
2L	280 m/sec

The rise velocity of an individual bubble is dependent upon the relative depletion level of the bubble, its geometry, and upon interactions with other plasma structure nearby. These first two effects are treated in Ossakow and Chaturvedi [1978], and the present results above are consistent with the results therein. For instance, the relatively high



rise velocity associated with 2S is seen to be due to the fact that the bubble is more severely depleted than that in 1S. Further, the roughly equal rise velocities of the 1S, 1L, and 2L bubbles, in spite of the fact that the 1L and 2L bubbles are much more severely depleted than that in 1S, is explained by noting that 1S actually approximates the geometry of a "sheet" of depleted plasma, whereas the 1L and 2L bubbles more closely resemble a cylindrical geometry (see Ossakow and Chaturvedi [1978]).

In an attempt to understand the reasons for the differences in the nonlinear evolution of small and large horizontal scale perturbations, we look at the potential equation:

$$\nabla^2 \phi_1 + \frac{\nabla(v_{in} n)}{v_{in} n} \cdot \nabla \phi_1 = \frac{-Bg}{c v_{in}} \frac{1}{n} \frac{\partial n}{\partial x}$$

At early times we expect  $\nabla \phi_1$  to be small with respect to  $Bg/cv_{in}$ , so we ignore the second term on the left hand side, giving a Poisson equation for  $\phi_1$ :

$$\nabla^2 \phi_1 = - \frac{Bg}{c v_{in}} \frac{1}{n} \frac{\partial n}{\partial x}$$

We can now interpret the right hand side as simply the local charge density  $\rho$  (such that  $\nabla \cdot \underline{E} = \rho$ ). Since we have initialized all of these calculations with  $\frac{1}{n} \frac{\partial n}{\partial x}$  independent of  $y$  (see (12)), what we are dealing with is a distribution of charge density that has the form of diffuse "plates" aligned in the vertical direction. Noting that the

term  $v_{in}$  decreases with altitude, we find that these diffuse "plates" have an equally diffuse "edge" in the  $y$  direction. Taking the analogy to its conclusion, we model our initial conditions, or any vertically aligned structure for that matter, as an array of plates of charge (non-conducting capacitor plates) with an edge somewhere in or above our grid. In Figure 8 we show schematically the electric potential field surrounding the edge of parallel plates of charge for two different separation distances. Since there is only one scale length in the configuration (the plate separation distance  $d$ ), then all other scale lengths must be proportional to this distance. In particular, the characteristic distance parallel to the plates over which the electric field outside the plates (the fringe field) falls off must be proportional to  $d$ . Since in our problem the contours of electrostatic potential are in fact streamlines (see (11)), this distance will determine the maximum depth in the fluid from which the electrostatic field will draw fluid into the bubble. Since the plasma density is lower at greater depths, this distance will determine the minimum plasma density inside the bubble. To test these ideas, we examine the actual early time electrostatic potential fields from the calculations we have presented. Fig. 9 shows contours of  $n$  and  $\phi_1$  for the 1S initial conditions, and the same plots for 1L. A similar comparison for cases 2S and 2L is shown in Fig. 10. All contour plots of  $\phi_1$  are scaled in such a way as to evenly space exactly 12 contour lines between the maximum and minimum value of  $\phi_1$ , to normalize the plots so they can be compared without bias. The increased vertical extent of the contours of  $\phi_1$  (streamlines) for

cases 1L and 2L are evident.

Of course, the initial profile generating these aligned plates of charge lasts only a short time. In the linear phase of growth, the perturbation grows in the region of linear instability (the F region bottomside), and damps elsewhere. Our "plates" therefore very quickly become horizontally spaced regions of charge with a limited vertical extent, confined to the region of steepest vertical gradient on the F region bottomside. Nonetheless, the scaling arguments advanced above still hold: the vertical extent of the polarization electric field scales as the horizontal extent of the structure causing the field. This is easily seen in Figure 11 and 12 where comparison is made of the  $\phi_1$  contours for 1S vs 1L and 2S vs 2L respectively, at a time of 700 sec. Cases 1S and 2S are seen to be mixing fluid over a fairly narrow altitude range, while 1L and 2L have each formed a large convective cell more than 150km in vertical extent, drawing plasma fluid into the bubble from deep in the ionosphere. It is not surprising, then, that the larger horizontal scale bubbles are more severely depleted at late times.

#### IV. SUMMARY

On the basis of our numerical simulations, and of a qualitative scale analysis of the driving electrostatic potential equation, we conclude that the severe "biteouts" of three orders of magnitude and bubble rise velocities of 150 m/sec reported by McClure et al [1977] are completely consistent with large (~75-200km) horizontal bubble

size scales. In our simulations, the severe biteouts associated with large horizontal scale lengths are due to the fact that the plasma comprising these bubbles has its origin at much lower altitudes than in the small horizontal length scale cases. Again, these results are consistent with those of McClure et al. [1977], who base their conclusions on ion mass spectrometer measurements of the  $H^+ - O^+ - N^+$  balance inside the bubbles, which they find to be "characteristic of undisturbed plasma found at lower altitudes." The variation in the vertical velocities of various bubbles noted by McClure et al. [1977] is probably due to interactions between bubbles. Note, for example, that in Fig. 2, the secondary bubble is rising at a much slower rate than is the central bubble. Bubble interaction will be the subject of forthcoming theoretical and numerical studies.

#### APPENDIX: Numerical Solution of the Equations

Of the two partial differential equations we must solve, (10) is elliptic and linear and (11) is hyperbolic and nonlinear. Both equations represent numerical challenges, and we could easily devote the bulk of this paper to the numerical techniques used for their solution. However, as we stated in the introduction, we shall confine ourselves to the improvements made in these techniques since the calculations of Ossakow et al. [1979]. We begin with (11).

Equation (11) is solved in finite difference form using a fully multidimensional second order in time, fourth order in space, leapfrog-trapezoidal flux-corrected transport (FCT) scheme. Both the higher order leapfrog-trapezoidal scheme itself, as well as the fully multidimensional algorithm utilized in the critical flux-limiting stage of FCT, are recent developments and are described by Zalesak [1979]. These developments represent significant extensions of the theory of FCT, a numerical technique originated by Boris and Book [1973] to handle equations of the form (11) where steep gradients are expected for form. By contrast, the calculations in Ossakow et al. [1979] used a first order in time, second order in space FCT algorithm which was only one-dimensional (since fully multidimensional FCT algorithms did not exist at the time), and hence used time-splitting (sequential x and y operators) to solve the two-dimensional equation (11). It is known that time-splitting can introduce numerical

problems into an incompressible flow calculation (see Zalesak, 1979), although our previous equatorial spread F calculations did not exhibit any of the symptoms of these difficulties.

Equation (10) is the elliptic equation whose solution gives us the electrostatic potential  $\phi_1$ . The right hand side is known and the left hand side represents a Hermitian operator operating on  $\phi_1$ , giving only real eigenvalues and apparently easing the difficulty of solution. However, the extremely large range for the values of the quantity  $v_{in} n$  makes for an equally large span of operator eigenvalues, and solution of the equation in this form using iterative techniques has not been successful. We have found one (and only one) method of direct solution, the stabilized error vector propagation (SEVP) method of Madala [1978], but the execution speeds for SEVP are not as favorable as for the method we now describe.

We start by expanding the operator and dividing through by  $v_{in} n$ , as was done by Ossakow et al. [1979], giving

$$\nabla_1^2 \phi_1 + \frac{\nabla(v_{in} n)}{v_{in} n} \cdot \nabla_1 \phi_1 = - \frac{Bg}{c} \frac{1}{v_{in} n} \frac{\partial n}{\partial x} \quad (A-1)$$

The equation is now in a form suitable for solution by the Chebychev-iterative relaxation technique of McDonald [1977]. However, great care must be given to the differencing of the term  $\frac{1}{v_{in} n} \nabla(v_{in} n)$  and  $\frac{1}{n} \frac{\partial n}{\partial x}$ , and this is the point we wish to address. We work with the term  $\frac{1}{n} \frac{\partial n}{\partial x}$  in one spatial dimension, since this example will make the point. A straightforward second-order difference form for this term is

$$\left(\frac{1}{n} \frac{\partial n}{\partial x}\right)_i = \frac{1}{2\Delta x} \frac{n_{i+1}^{-n_{i-1}-1}}{n_i} \quad (\text{A-2})$$

where the subscript  $i$  refers to grid point location in the  $x$  direction and  $\Delta x$  is the (uniform) spacing between grid points. This is the form used in Ossakow et al. [1979]. We shall show that this difference form produces solutions  $\phi_1$  with potentially undesirable properties, and causes undue numerical hardship in finding those solutions.

Let us rewrite  $\frac{1}{n} \frac{\partial n}{\partial x}$  as  $\frac{\partial}{\partial x} (\ln n)$ . A physical interpretation of the term is now much easier: the term tells us how rapidly the logarithm of  $n$  is varying with respect to  $x$ . Suppose, for argument's sake, that the smallest and largest values of  $n$  in the problem are  $10^1$  and  $10^5$  respectively. On a grid of size  $\Delta x$ , the largest value representable for that term would occur when a fluid element of density  $10^1$  and one of density  $10^5$  occupy adjacent grid points. The value of  $\frac{\partial}{\partial x} (\ln n)$  evaluated midway between these two grid points would be  $\frac{1}{\Delta x} (\ln 10^5 - \ln 10^1) = 9.2 \Delta x^{-1}$ . Evaluation of (A-2) for  $n_{i-1}, n_i$ , and  $n_{i+1}$  having values of  $10^1, 10^1$ , and  $10^5$  respectively gives a value for  $\left(\frac{1}{n} \frac{\partial n}{\partial x}\right)_i$  of  $5 \times 10^4 \Delta x^{-1}$ , far in excess of the maximum value for this term given by the above argument. Logarithmic interpretation of this term would state that  $n$  varied by more than  $10^4$  orders of magnitude over a single grid spacing, a ridiculous statement in light of the fact that there are only four orders of magnitude of  $n$  in the problem.

The potentially large values of these terms, in particular of the term  $\frac{1}{v_{in} n} \nabla_{\perp} (v_{in} n)$  in (A-1), not only cause extremely slow convergence of the iterative solution, but can also put **spurious** oscillations into the exact finite difference solution  $\phi_1$  due to cell Reynolds number effects [Roache, 1976]. As shown by Roache [1976] these oscillations can occur any time the value of the term  $\frac{1}{v_{in} n} \nabla_{\perp} (v_{in} n)$  in (A-1) exceeds a critical value of  $2 \Delta x^{-1}$ . It is clear, then, that (A-2) represents an undesirable difference form for these logarithmic terms. Better approximations are

$$\left( \frac{1}{n} \frac{\partial n}{\partial z} \right)_i = \frac{1}{2\Delta x} (\ln n_{i+1} - \ln n_{i-1}) \quad (A-3)$$

and 
$$\left( \frac{1}{n} \frac{\partial n}{\partial x} \right)_i = \frac{1}{\Delta x} \frac{n_{i+1} - n_{i-1}}{n_{i+1} + n_{i-1}} \quad (A-4)$$

Equation (A-3) is probably the most accurate, but evaluation of the logarithms at every time step is computationally expensive, and for problems like the one at hand where  $n$  varies by many orders of magnitude across the grid, there is still no guarantee that the critical cell Reynolds number will not be exceeded. For these reasons we use (A-4) for the present calculations.



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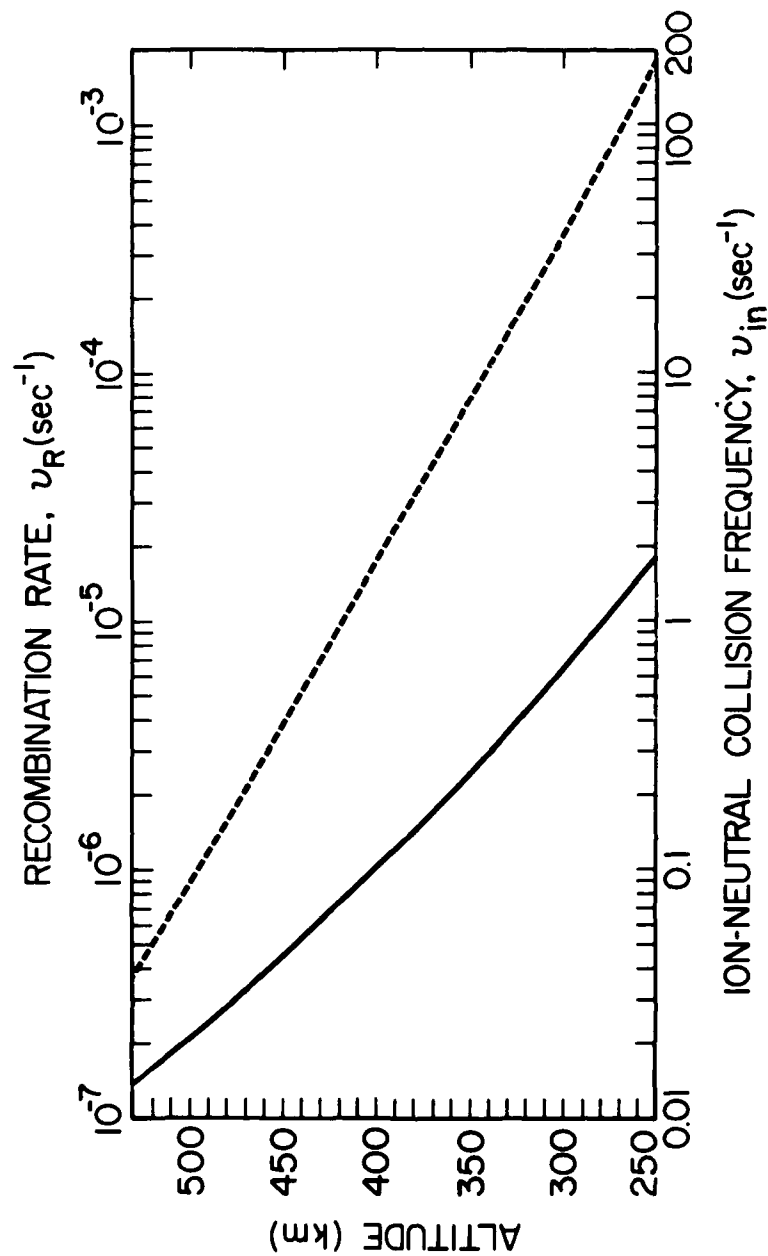


Fig. 1 - The ion-neutral collision frequency  $\nu_{in}$  and recombination rate  $\nu_R$  (as a function of altitude) used in the numerical simulations

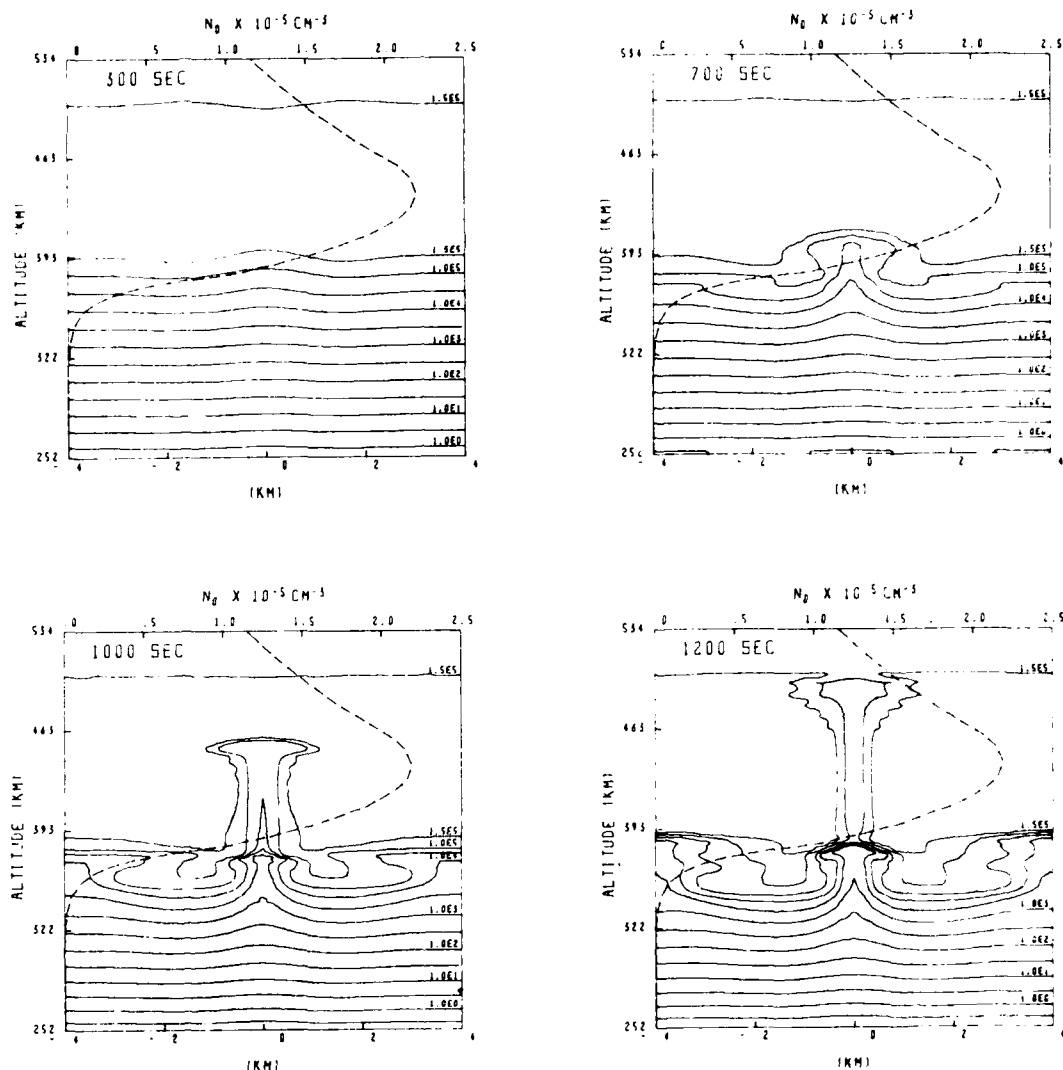
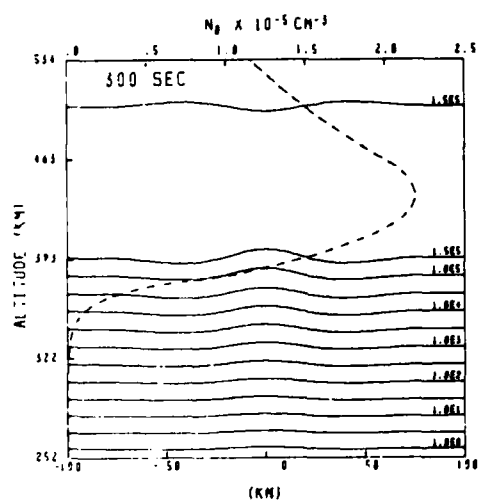


Fig. 2 - Sequence of four plots showing iso-electron density contours of calculation 1S at 300, 700, 1000, and 1200 sec. Superimposed on each plot is a dashed line depicting  $n_0(z,t)$ . Electron densities are given in  $\text{cm}^{-3}$ . The observer is looking southward.



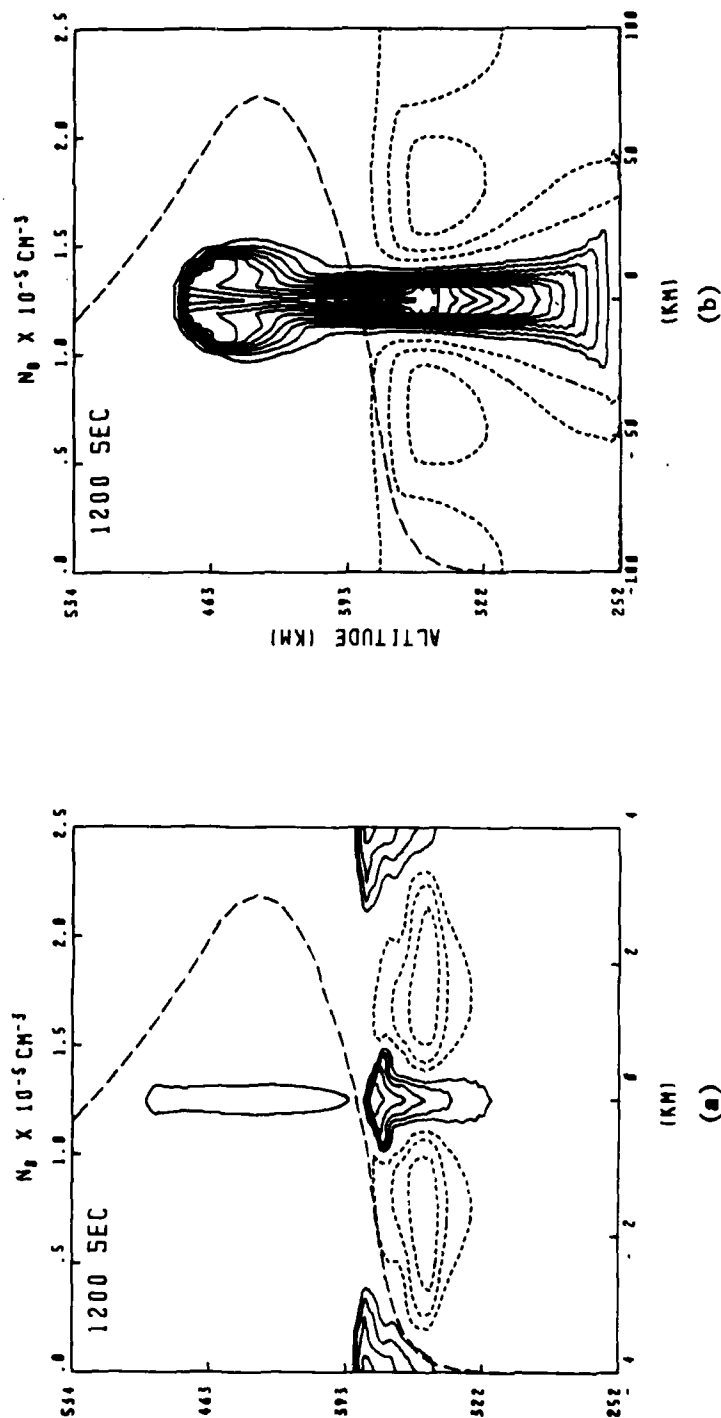
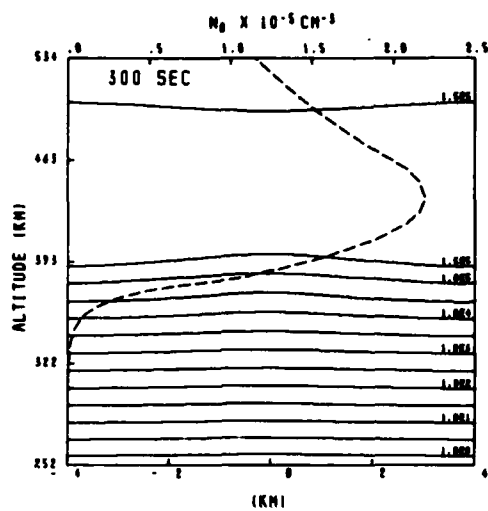
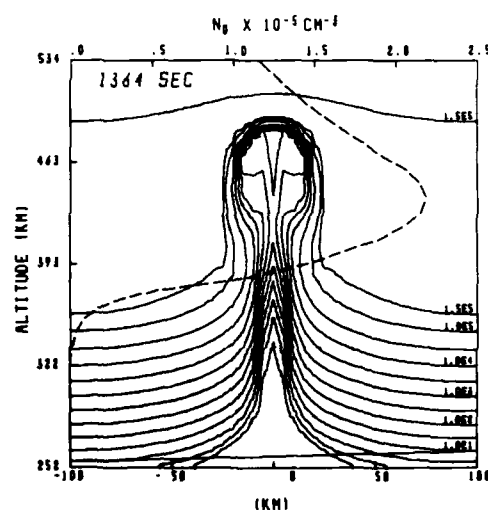
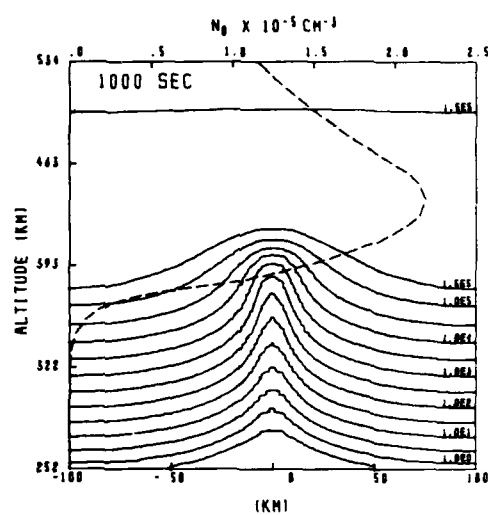


Fig. 4 - Contours of constant  $n(x,y,t)/n_0(z,t)$  for a) 1S at 1200 sec and b) 1L at 1200 sec. Depletions ( $n/n_0 < 1$ ) are shown in solid lines while enhancements ( $n/n_0 > 1$ ) are shown as short dashed lines. The first (outermost) depletion contour is for  $n/n_0 = 0.5$ , while each succeeding contour is for a value of  $n/n_0$  a factor of 0.5 times the previous one. The first enhancement contour is for  $n/n_0 = 2.0$ , while each succeeding contour is for a value of  $n/n_0$  a factor of 2.0 times the previous one. The superimposed long dashed line depicts  $n_0(z,t)$ .







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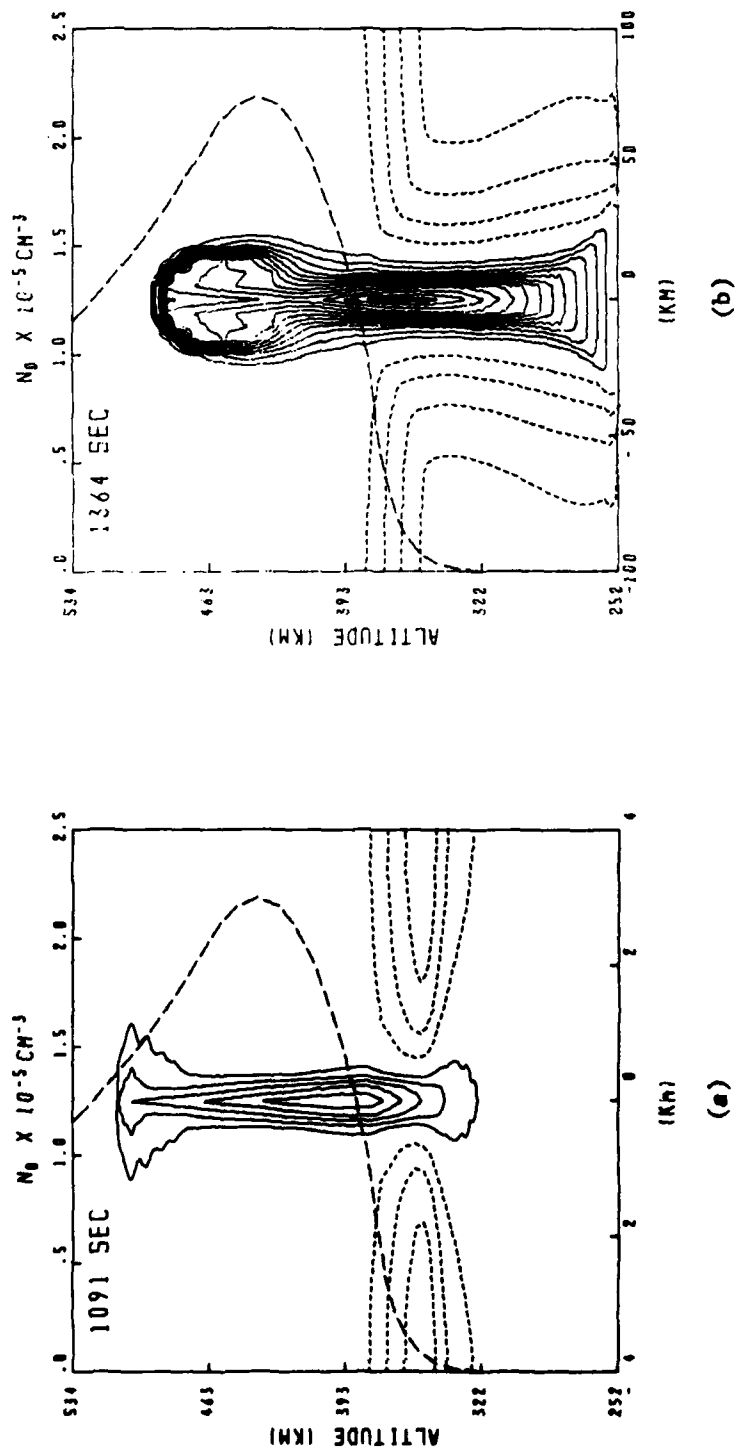


Fig. 7 - Same as Fig. 4 but for a) calculation 2S at 1091 sec and  
b) calculation 2L at 1364 sec.

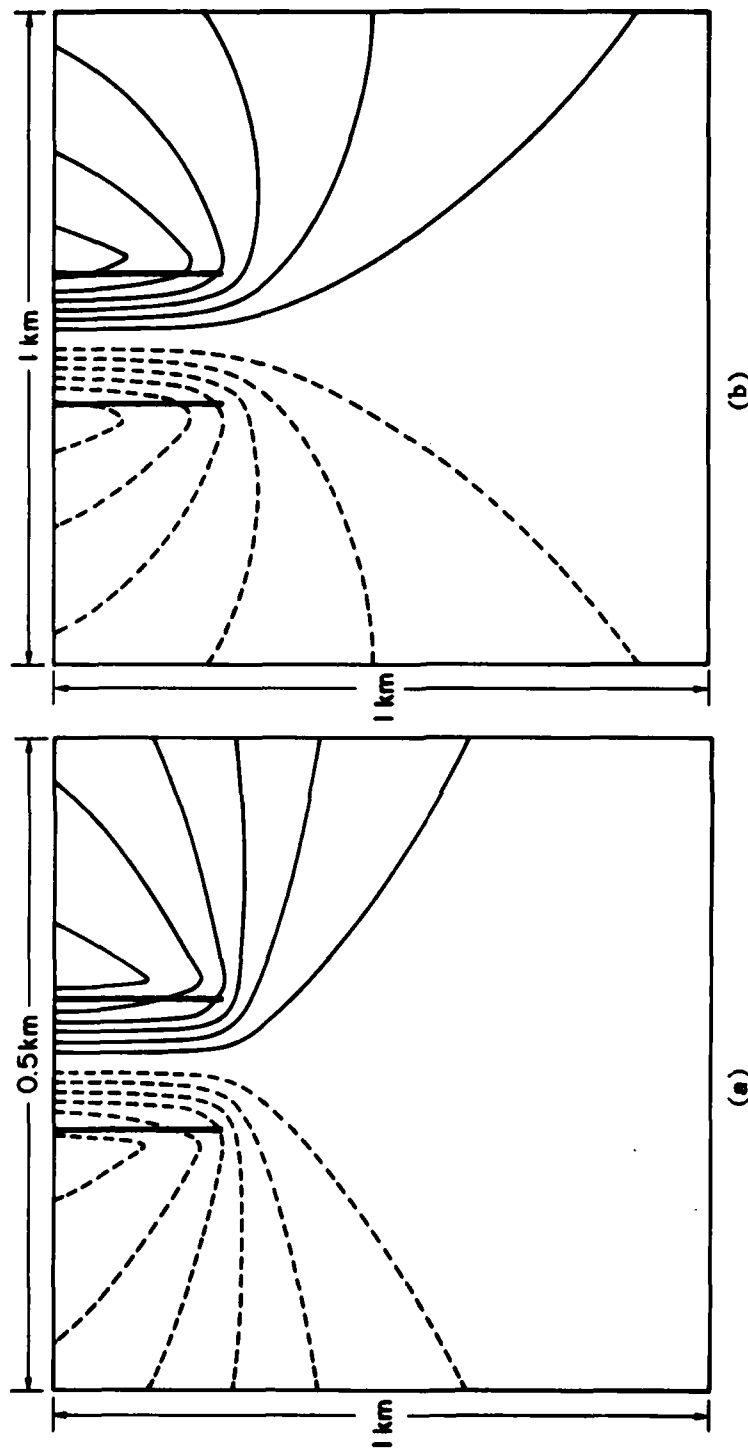


Fig. 8 - Schematic representation of contours of constant electrostatic potential produced at the edge of non-conducting parallel plates of charge for two different plate separation distances: a) 0.5 km and b) 1.0 km. Dashed lines are for negative potentials and solid lines are for positive potentials. The potential drop across adjacent contours is a constant in each plot, except that the zero potential (which would just be a vertical line) is not plotted. In each case the contours are chosen to space the twelve contours (plus the zero contour) uniformly from maximum to minimum potential. Since the electric field magnitude is inversely proportional to the contour spacing, we have, in effect, normalized each plot to the maximum electric field found between the plates. This facilitates the comparison of the rate at which the electric field falls off away from the edge of the plates. The much more rapid falloff in a) is evident.

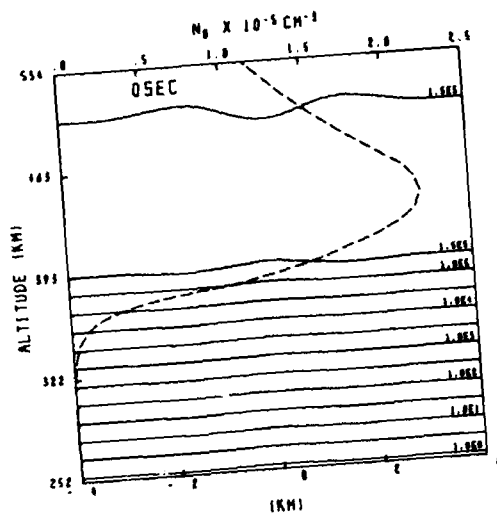


Figure 9(a)

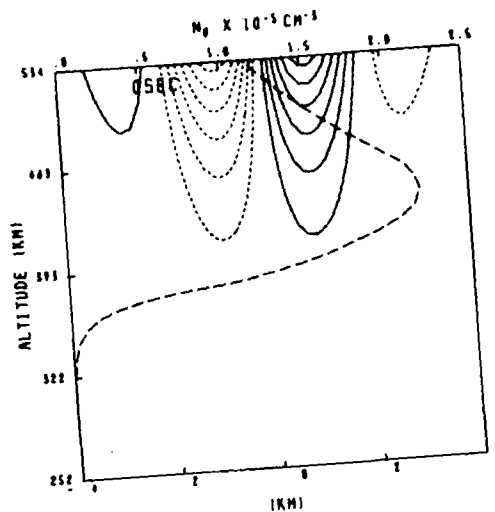


Figure 9(b)

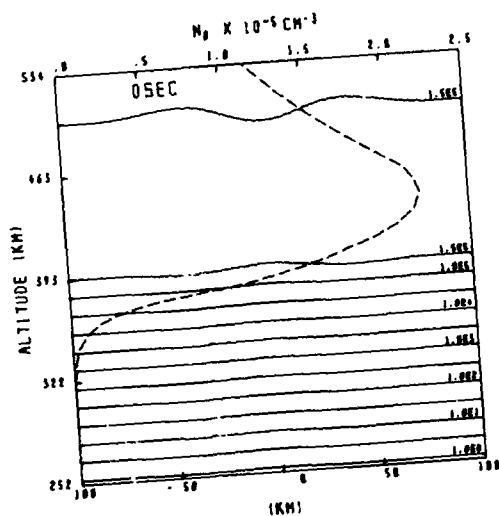


Figure 9(c)

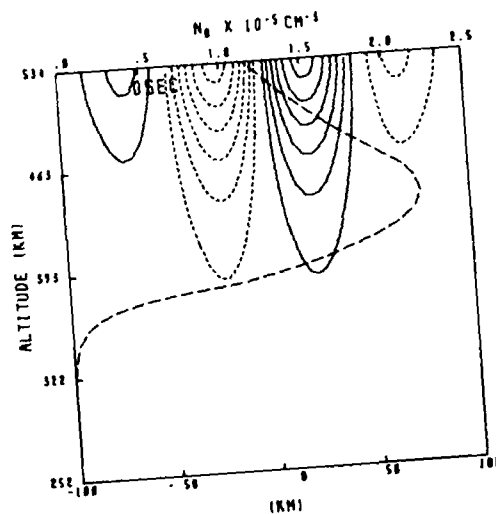


Figure 9(d)

Fig. 9 - Iso-electron density plots at  $t = 0$  for 1S (a) and 1L(c). Also shown are the corresponding contours of constant induced electrostatic potential in (b) and (d) respectively. For the plots of electrostatic potential, the comments from Fig. 8 apply. Since the contours of constant potential are in fact flow streamlines, more plasma from lower altitudes is being drawn upward in the 1L case than in the 1S case.



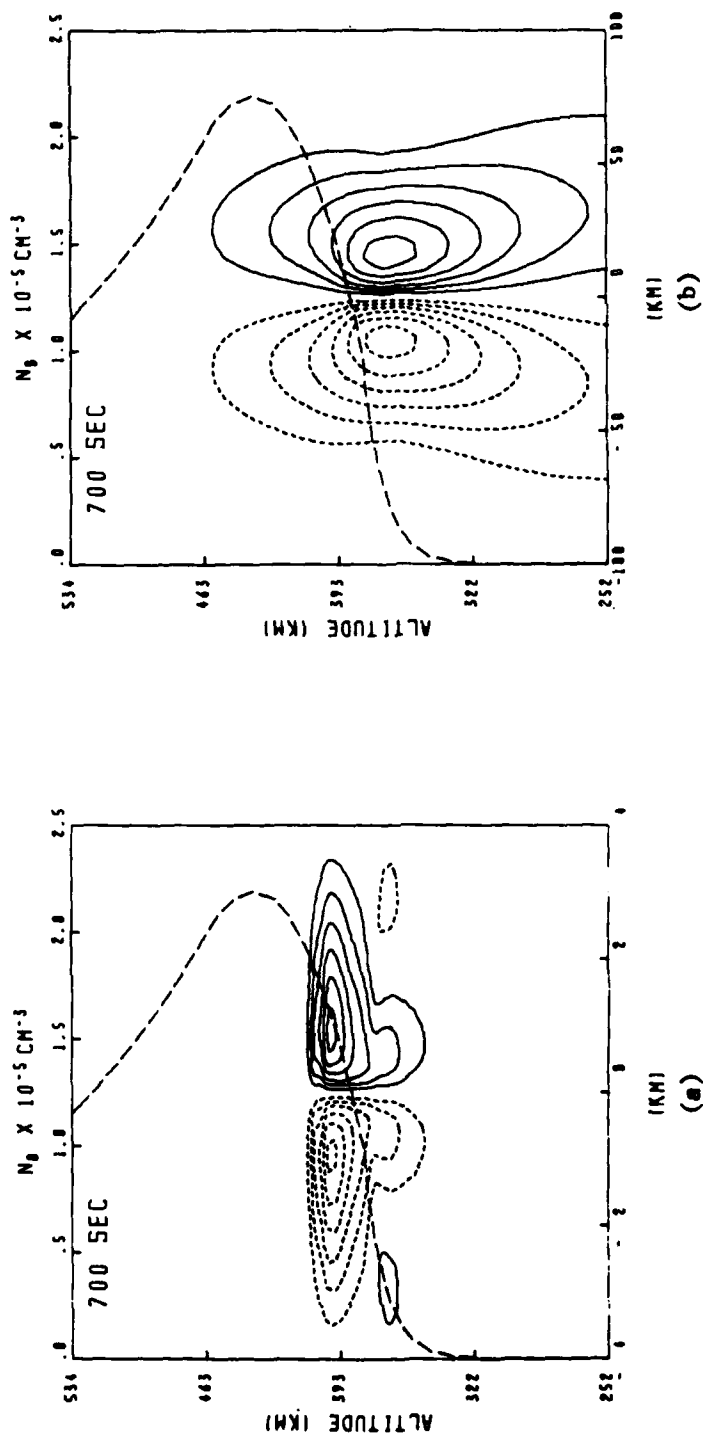


Fig. 11 - Contours of constant electrostatic potential at 700 sec for (a) calculation 1S and (b) calculation 1L. Comments from Fig. 8 and Fig. 9 apply.

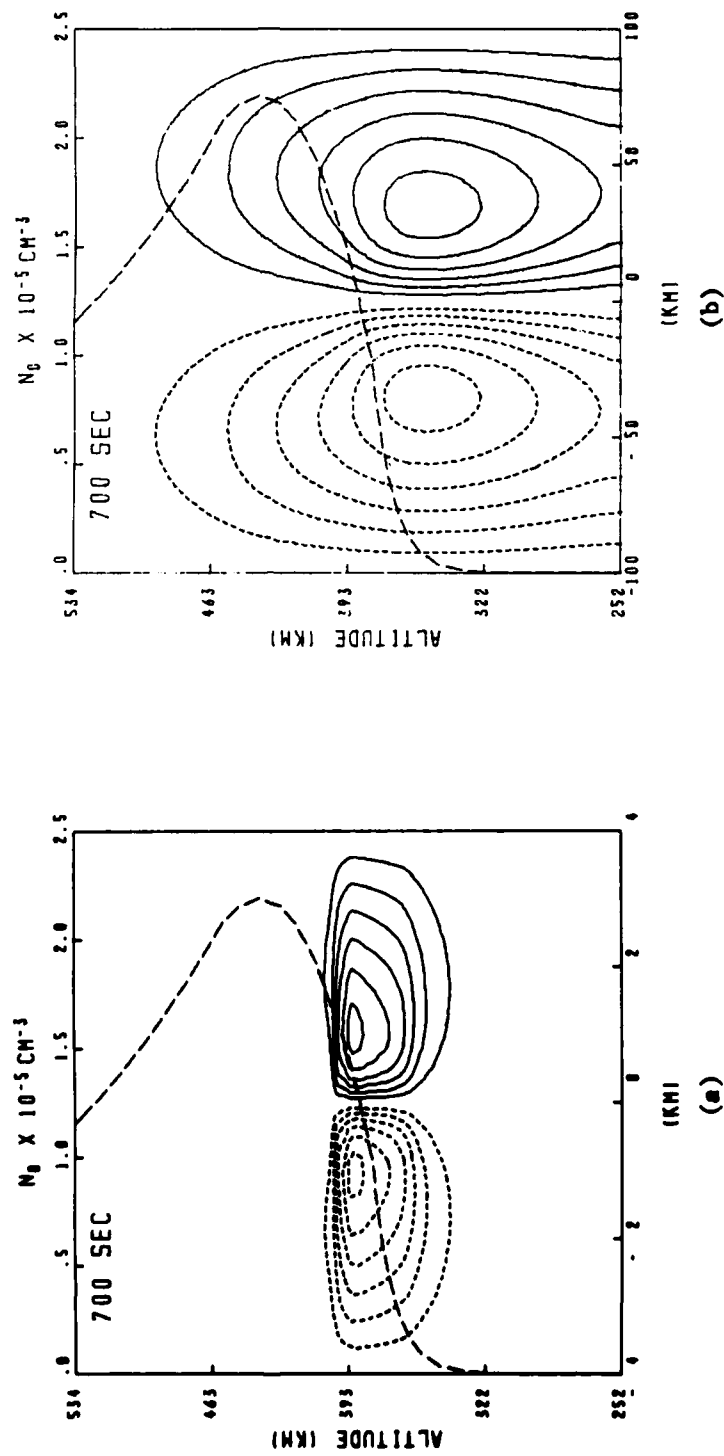


Fig. 12 - Same as Fig. 11 but for (a) calculation 2S and (b) calculation 2L.

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